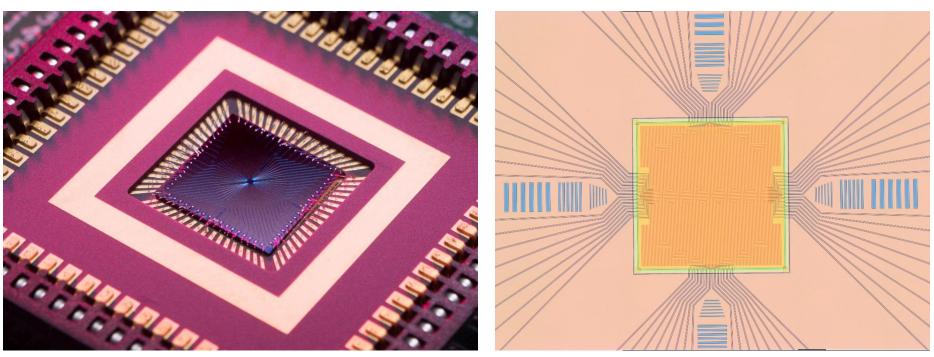
Superconducting Nanowire Single Photon Detectors for Optical Communication and Quantum Optics

Matt Shaw

APh 110 Seminar, 2 October 2017

Jet Propulsion Laboratory, California Institute of Technology



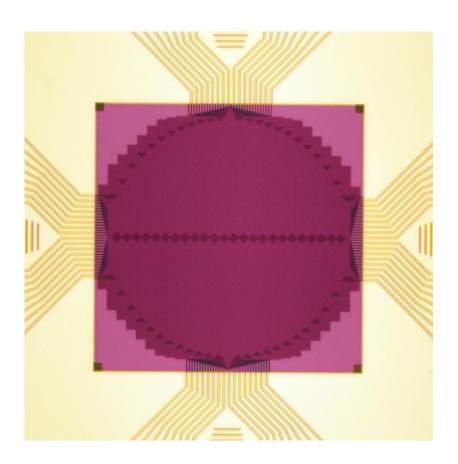
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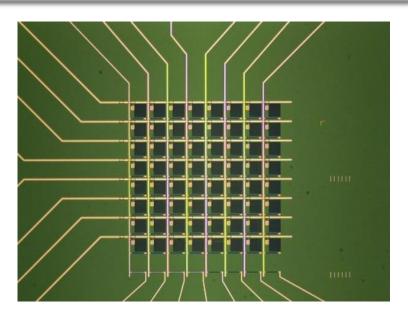


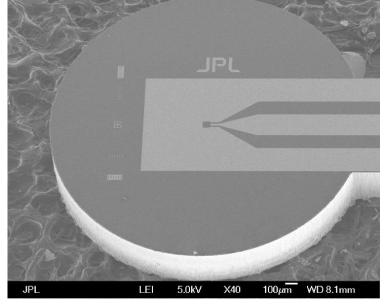
Research Group Overview

We develop **Superconducting Technology** to solve problems in single-photon detection

We use this technology to enable new science









JPL SNSPD Development Team

JPL Staff



Matt Shaw



Andrew Beyer



Ryan Briggs



Emma Wollman



Marc Runyan



Angel Velasco

Huy Nguyen

Postdocs

Graduate Students



Boris Korzh



Jason Allmaras

Alumni



Jeff Stern 1962-2013



Francesco Marsili



Bill Farr

Visiting Students



Eric Bersin



Simone Frasca

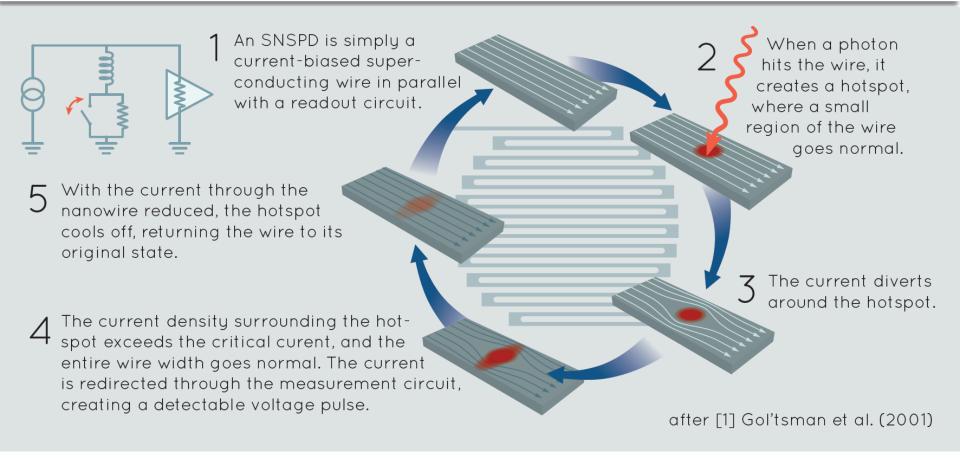


Eddy Ramirez

Kelly Cantwell Chantel Flores Sarang Mittal Marco Suriano Luca Marsiglio Giovanni Resta Garrison Crouch Andrew Dane Emerson Viera Viera Crosignani Michael Mancinelli Neelay Fruitwala



SNSPD Device Concept



- Performance exceeds conventional detectors but requires cryogenic cooling (1-4 K)
- 16 years of active development (MIT / LL, NIST, JPL, Russia, Europe, Japan, China)
- SNSPDs are the highest performing detector for time correlated single photon counting

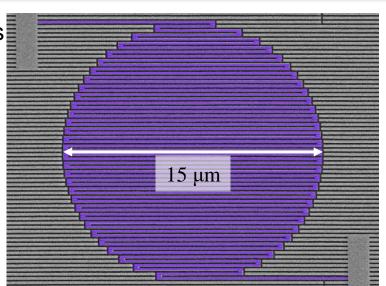


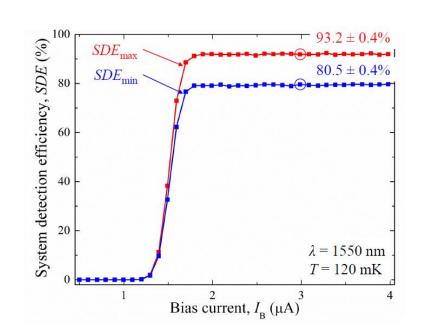
Tungsten Silicide SNSPDs

Superconducting Nanowire Single Photon Detectors

- Time correlated single photon counting, UV-5 μm
- > 90% efficiency at 1.5 μm
- < 100 ps time resolution
- Sub-Hz intrinsic dark rates
- 40 ns dead time
- Arrays as large as 64 pixels
- 1 K operating temperature
- Primary applications are space-based optical communication and quantum optics

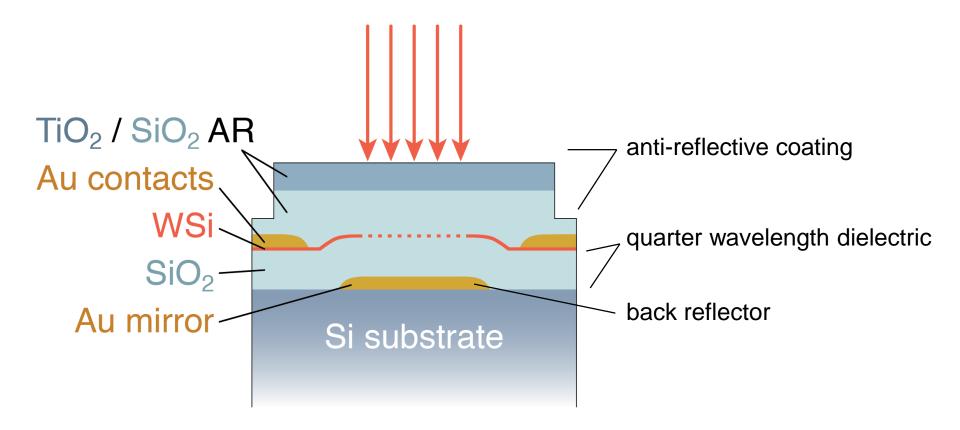
Marsili et al, *Nature Photonics* **7**, 210 (2013)







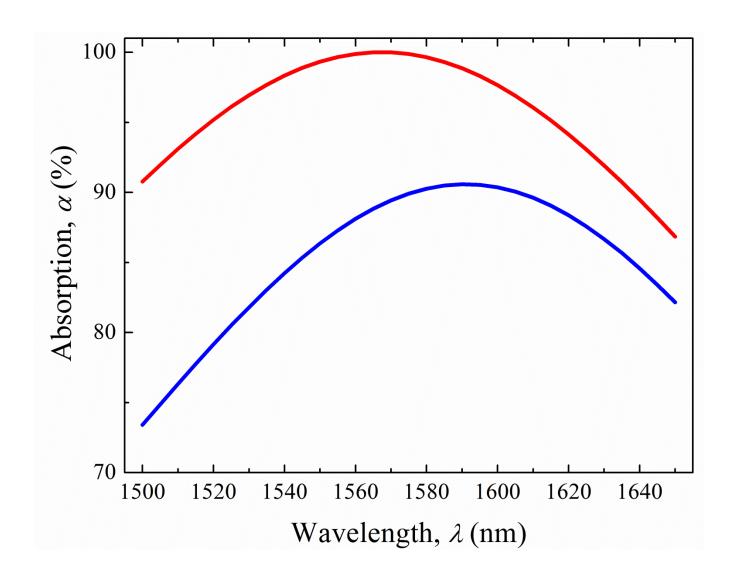
Optical Cavity to Enhance Absorption



- Photosensitive nanowire element is embedded in a vertical quarter-wave cavity
- Enhances detection efficiency from 20% to >90%
- Have optimized for high efficiency at 1550, 1310, 800, 373, 313 nm

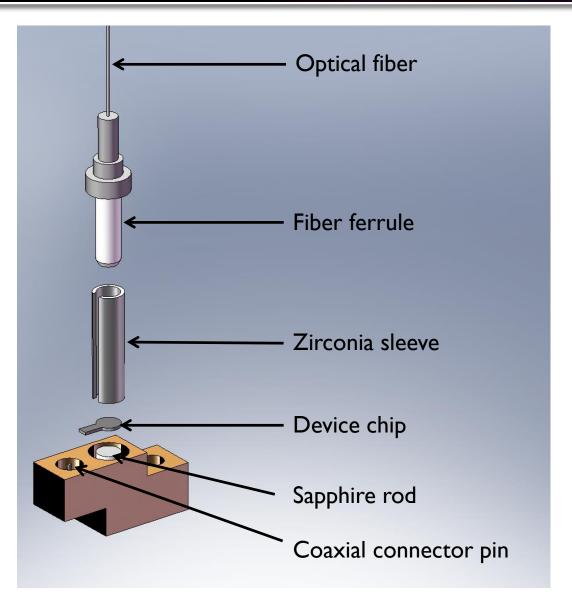


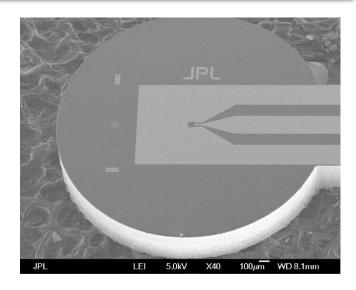
Optical Stack for Enhanced Absorption

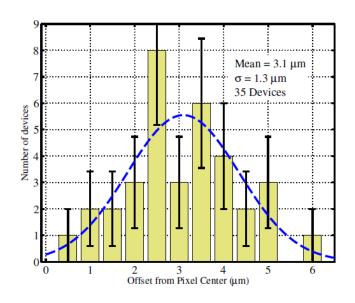




Self-aligned fiber coupling

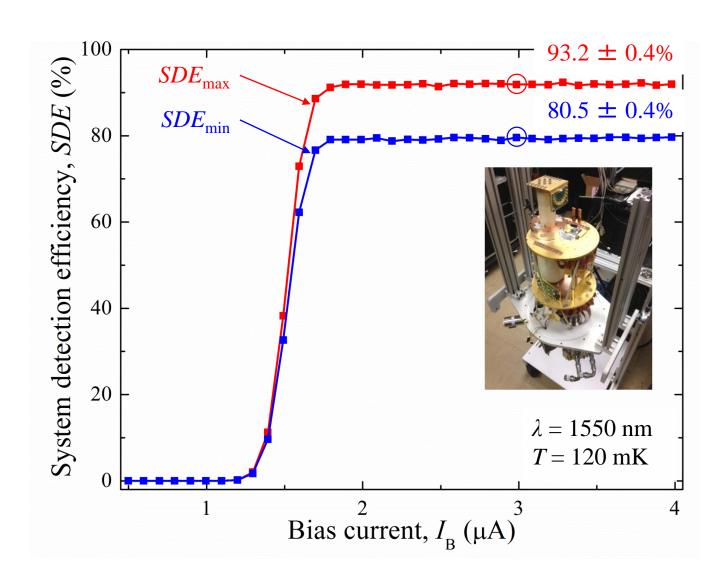






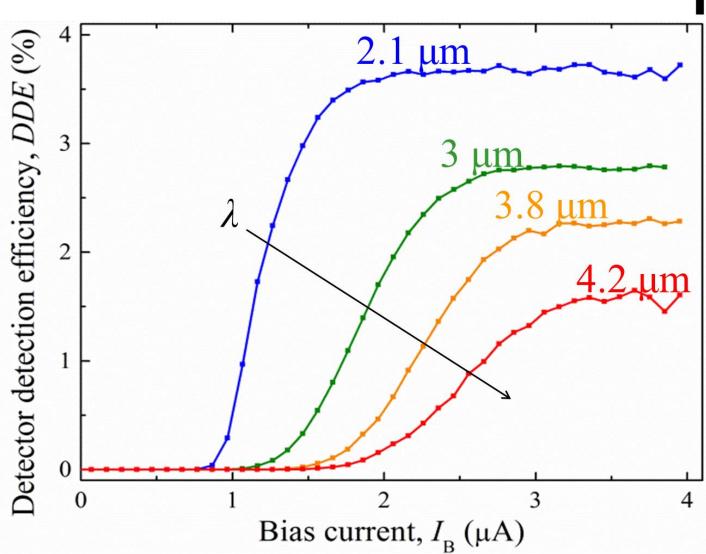


93% System Detection Efficiency





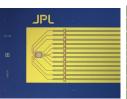
NST

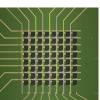


IR Arrays for Optical Communication

- 64 pixel arrays for DSOC ground terminal
- 12 pixel arrays for secondary LLCD ground terminal
- 64 pixel imaging arrays with row-column readout
- Fiber-coupled arrays for QKD
- Feasibility studies for future ISS teleportation concept







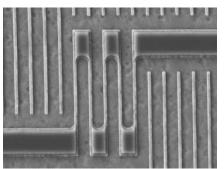


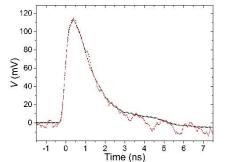
NASA STMD - NASA SCaN - DARPA

Other Government Agency

High Operating Temperature SNSPDs

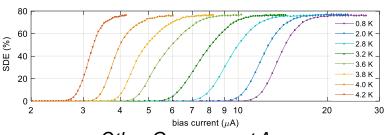
- MgB2 SNSPDs single photon sensitive at 17 K
- Currently exploring YBCO SNSPDs w/ HeNeGa FIB





Ultraviolet SNSPDs

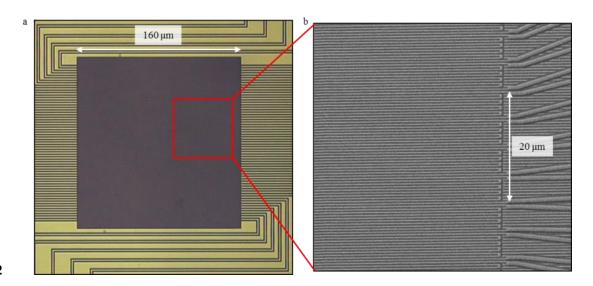
- UV SNSPDs for trapped ion quantum computing
- 80% efficiency at 370 and 315nm at 4.2 K
- Single photon sensitivity at 245nm
- Potential future applications in UV astrophysics

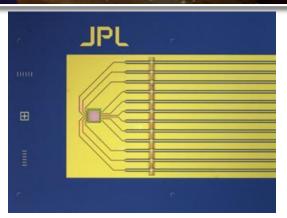


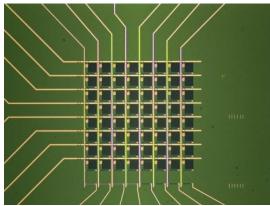
Fundamental Research

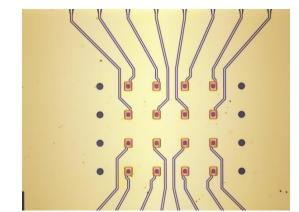
- Ultra-low jitter SNSPDs
- Frequency and time-domain multiplexing schemes
- Waveguide SNSPDs for integrated photonics
- Superconducting readout electronics
- Theory and device physics modeling
- New detector concepts and applications
- Collaboration with quantum optics groups

- Amorphous WSi permits scaling to large arrays
- Each pixel has essentially the performance of a single pixel fiber-coupled device
- Fiber-coupled vs free-space
- Non-imaging vs imaging
- Direct readout vs "row-column"
- Recently demonstrated up to 64 pixel arrays

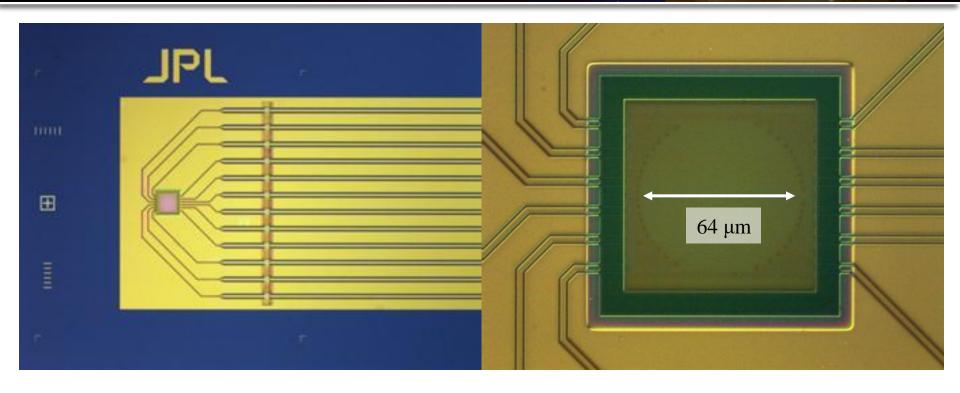








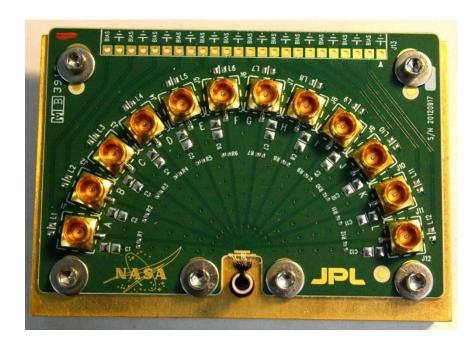


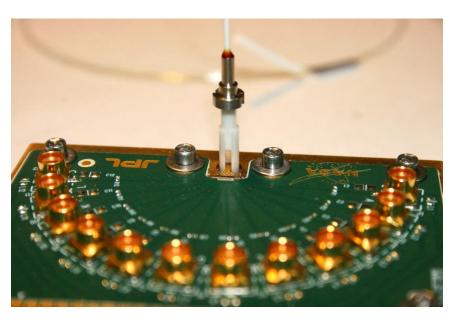


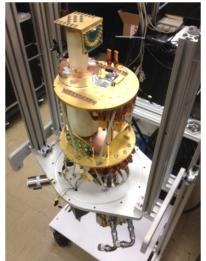
- Array design based on single-pixel SNSPDs described above
- 12 Counterbiased Nanowires, 6 wires in each plane, 160 nm width
- Designed for fiber coupling to 1-m OCTL telescope
- 64 µm diameter active area matched to GIF-625 multimode fiber
- Quarter wave optical cavity to resonantly enhance absorption at 1.55 µm
- ~40% total array efficiency
- Max count rate ~10 MHz / channel



12 Pixel SNSPD Array Packaging







- Same self-aligned fiber coupling scheme is used
- Each wire is read out independently
- Operated in an 800 mK cryostat with fiber coupling

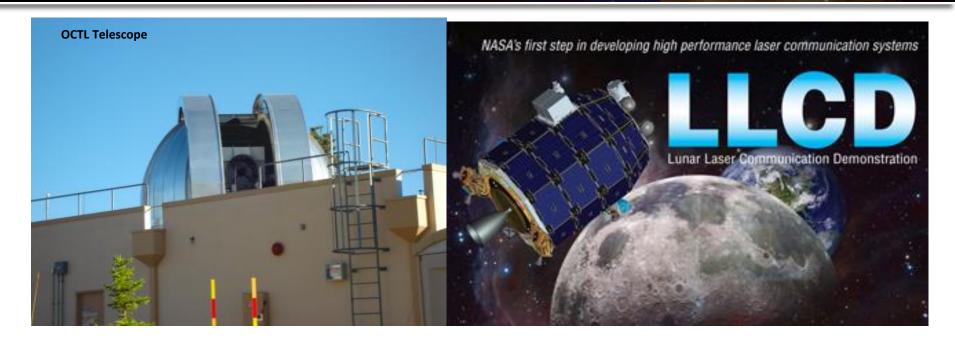


Field Demonstration: LLCD



- Bidirectional laser communication demo from lunar orbit (400,000 km) at 1550 nm
- First demonstration of laser communication beyond earth orbit
- Uplink rates 10-20 Mbps, Downlink rates 39-622 Mbps
- Transmit Payload on LADEE Spacecraft (ARC) implemented by MIT-LL
- Managed by GSFC, Primary ground terminal implemented by MIT-LL
- Secondary ground terminals implemented by JPL and ESA

LLOT Reciever Overview

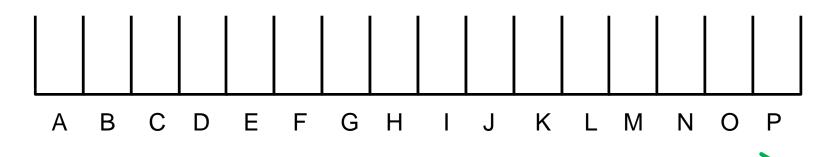


- LLOT was a secondary ground station for the Lunar Laser Communication Demonstration (LLCD), located at Table Mountain, CA
 - LLOT supported 20 LLST passes during October-November 2013
 - Received error-free downlink at 39 / 79 Mbps
 - Link support at Sun-Earth-Probe (SEP) > 10°
 - Transmit laser beacon to assist link acquisition
 - Transmit limited real-time channel and link diagnostics to operations center
 - Receiver implemented in software
 - 12-pixel WSi SNSPD detector arrays were operating at 800 mK.
 - SNSPD arrays were infused in ~January 2013, still very early stage technology



Pulse Position Modulation

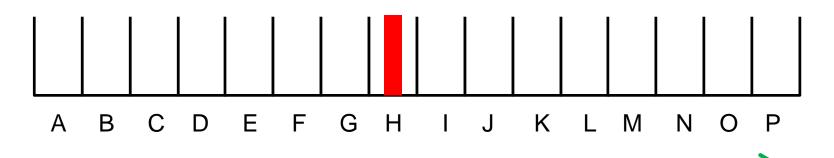
- Data is encoded in optical pulse timing
- Ideal for "photon starved" applications such as deep space optical communication
- Send more data with less mass and power on the spacecraft
- LLST signals use PPM-16 encoding with 311 MHz 5 GHz variable slot rate
- Scheme was recently used under DARPA InPho program to encode 13 bits/photon
- Requires detectors with high efficiency and sub-nanosecond time resolution





Pulse Position Modulation

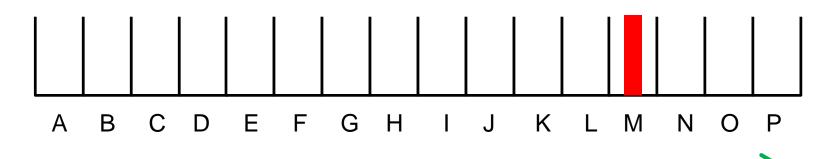
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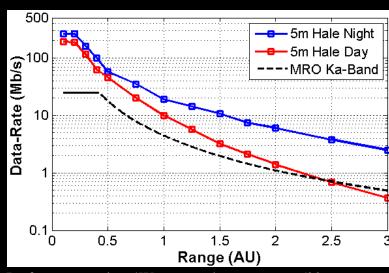


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DSOC Tech Demo Mission



Performance using 4W average laser power w/22 cm flight transceiver to 5m ground telescope

Beacon & Uplink 1030 nm 292 kb/s @ 0.4 AU

Spacecraft Flight Laser **Transceiver** (FLT) 4W, 22 cm dia

Ground Laser Transmitter (GLT) Table Mtn., CA 5kW, 1m-dia. Telescope

Ground Laser Receiver (GLR) Palomar Mtn., CA 5m-dia. Hale Telescope

Optical Comm Ops Ctr. JPL, Pasadena, CA

TBD MOC

Deep Space

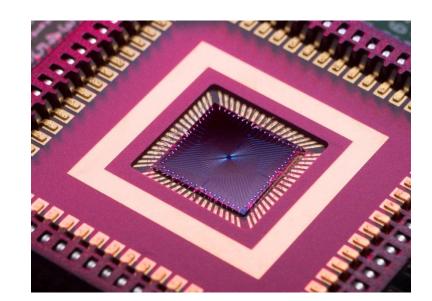
Network (DSN)



DSOC Project Overview

- Phase A of NASA Technology Demonstration Mission
- Flight terminal planned to launch on PSYCHE mission in 2022
- Projected downlink data rates from 200 kbps 265 Mbps
- PPM 16 128, 500 ps 8 ns slot widths, 4 slot intersymbol guard time
- Developing a 320-µm 64-pixel WSi SNSPD array for the ground receiver

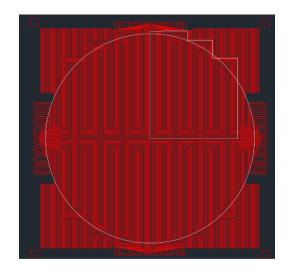




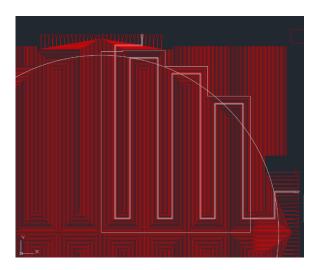


64-Pixel SNSPD Array

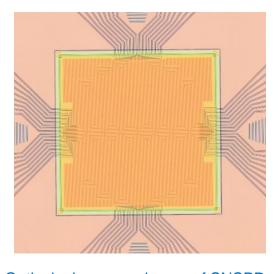
- 64-pixel WSi SNSPD array embedded in optical cavity optimized for 1550 nm
- 320-µm dia. free-space coupled active area, 4 quadrants, 16 co-wound wires per quadrant
- 13.3% nanowire fill factor: 4.5 x 160 nm wires on a 1.2 µm pitch
- Two-layer AR coating to enhance efficiency at low fill factor: 73% system detection efficiency
- 62 out of 64 measured nanowires show bias plateau
- Full 64-channel readout system and 64-channel time-to-digital converter



CAD Design of SNSPD focal plane array



CAD Design showing one of 16 individual sensor elements per quadrant



Optical microscope image of SNSPD array



Project Goals and SNSPD Performance

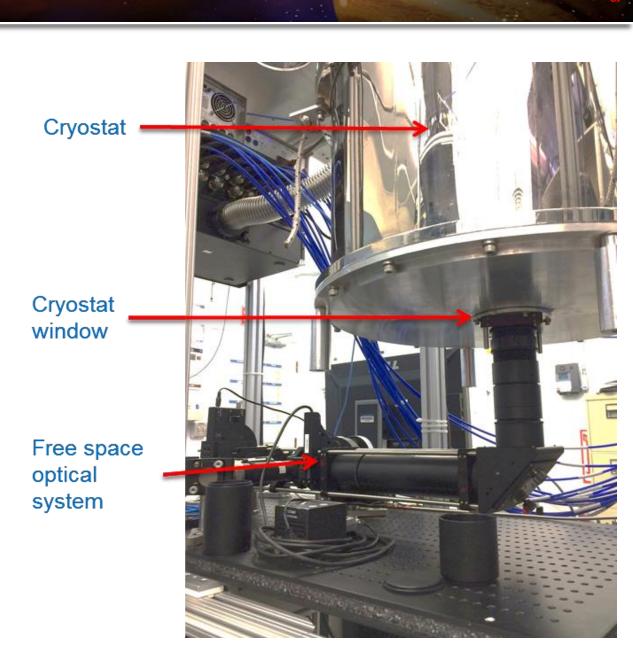
	DSOC Goals	Progress to date	Capability Achieved
Detection Efficiency	>50% minimum >70% desired	93% (fiber-coupled single pixel) 73% (Free Space, 320 µm array)	/
Timing Jitter	230 ps (FWHM)	80 - 120 ps (SNSPD array) 200 - 250 ps (with TDC)	
False Counts	< 1 Mcps total free space coupled	0.35 – 3.8 Mcps (320 μm array)	
Maximum Count Rate	830 Mcps (264 Mbps, 0.2 AU, night cruise)	465 – 1160 Mcps (SNSPD array) ~600 Mcps (with TDC)	
Active Area	260 µm diameter (35 µrad seeing, Palomar daytime)	320 µm diameter (50 µrad seeing, Palomar daytime)	
Numerical Aperture	f/1.2	f/4	

1550 nm operating wavelengthFree space coupled1 K operating temperature



Free Space Coupling

- Efficient coupling to large apertures requires free space coupling with cryogenic lens
- 300 K BK7 vacuum window
- 40 K, 4 K BK7 filters to block thermal background
- Engineering tradeoff between efficiency and false counts
- Experimenting with cryogenic spectral and spatial filters
- Must consider finite numerical aperture of detector



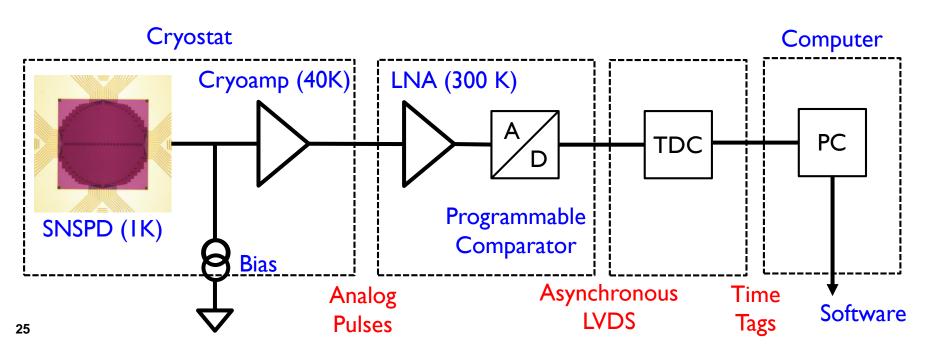


SNSPD Array Readout Architecture

- Direct readout of 64 channels into an FPGA
- Brass flex circuits from < 1 40 K
- DC-coupled cryogenic amplifiers
- Copper flex circuits from 40 300 K
- Room temperature amplifiers and comparators
- FPGA-based time tagger
- Currently setting up SNSPD optical communication testbed



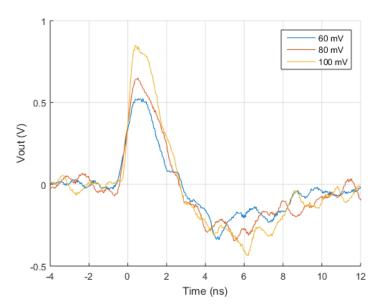
16-channel brass RF flex circuit



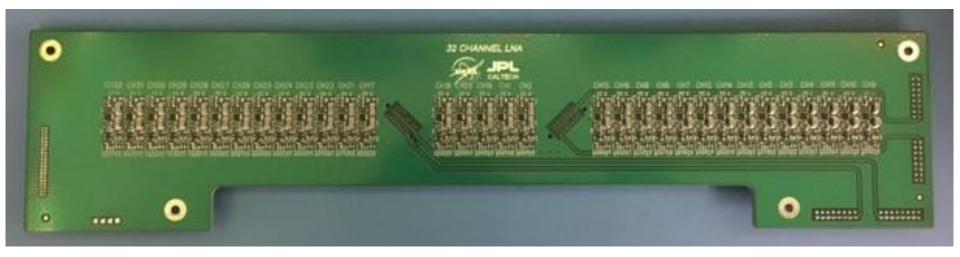


DC Coupled Cryogenic Amplifiers

- 2x 32-channel amplifier boards operated at 40 K
- 32 dB total gain
- Low-cost commercial cell phone components
- RFMD SGL-0622z cryogenic RF amplifier
- Broadcom ATF-35143 Psuedomorphic HEMT
- DC coupled with 50 ohm terminated input
- Detector bias added on amplifier board



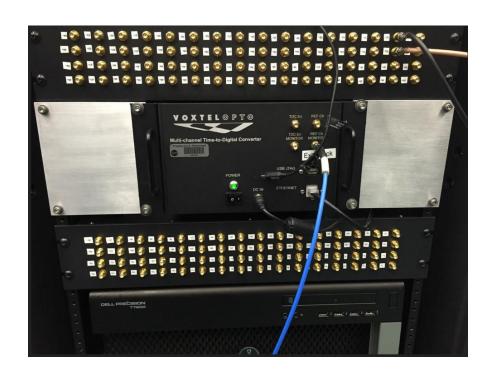
SNSPD output pulses at different bias points





Time to Digital Converter Development

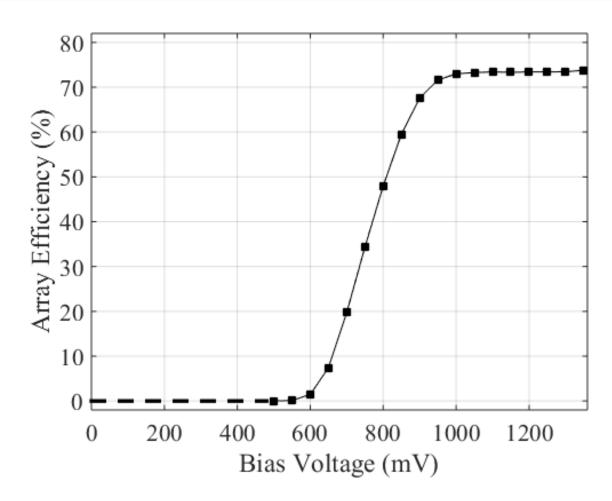
- Asynchronous time tagging receiver approach
- Need to tag photon arrivals across 64 channels with high time resolution
- Need to stream data into receiver FPGAs at ~ gigatag / second count rates
- Streaming 64-channel TDC currently under development through commercial partners
- Prototype TDC can fill 512 Mtag buffer at rates up to 600 Mtps w/ 166ps resolution







Efficiency Measurements

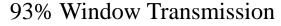


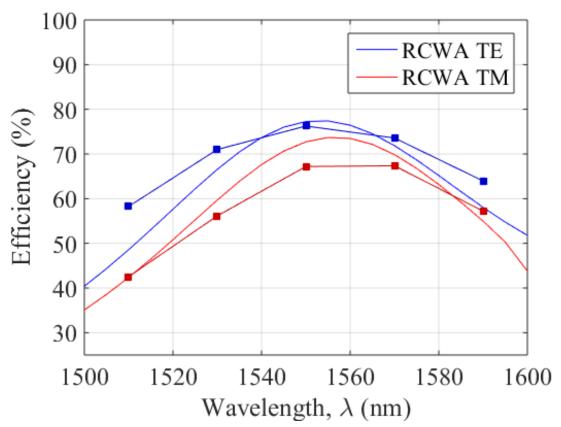
- System detection efficiency measured through cryostat window, 40K and 4K IR filters
- Measured SDE by focusing 50 um spot into one half plane (32 channels)
- Measured 73% efficiency in TE polarization at 1550 nm, 65% in TM



Efficiency Measurements

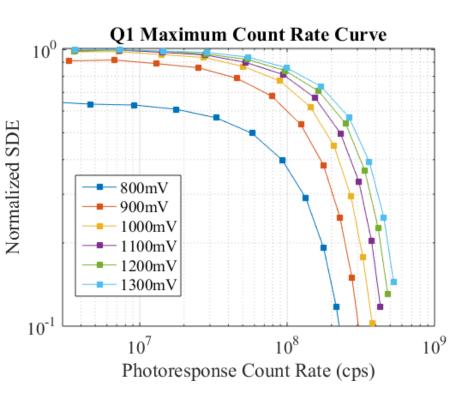
- Cavity is well centered near 1550 nm
- Efficiency matches RCWA simulation assuming 93% total transmission (97.6% per window)

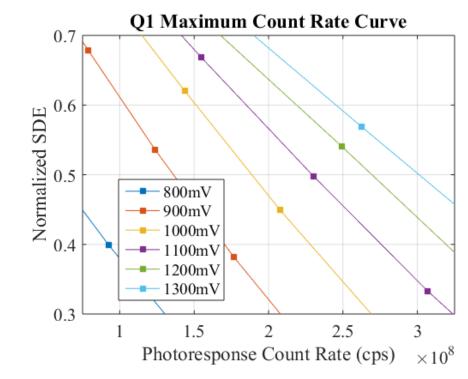






Maximum Count Rate



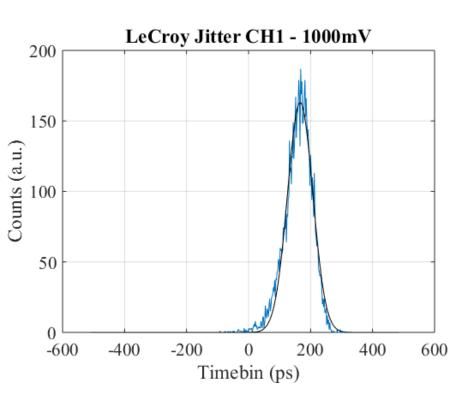


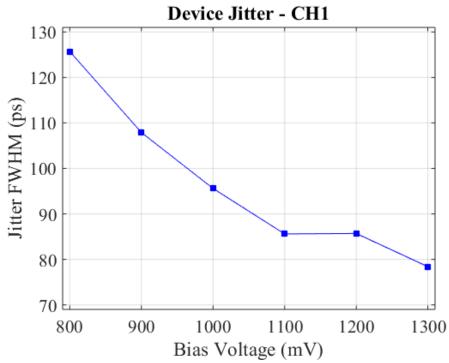
- MCR measured with beam centered on a single quadrant due to count rate limitations in TDC
- 120 300 Mcps 3dB point per quadrant
- Scales to 465 1160 Mcps across 62 pixels



Device Timing Jitter

- Representative individual pixel timing jitter measured using mode-locked laser and oscilloscope
- IRF is close to Gaussian
- 125 79 ps FWHM

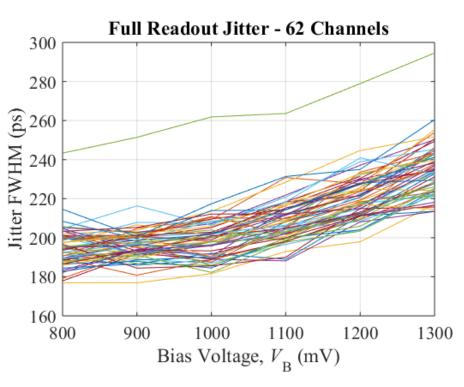


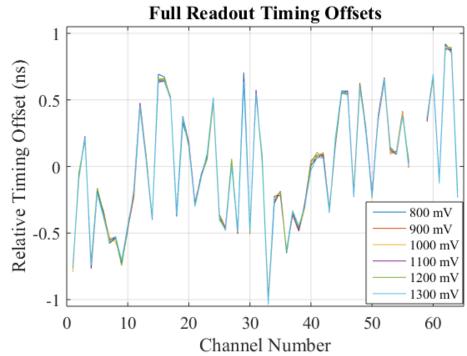




System Timing Jitter

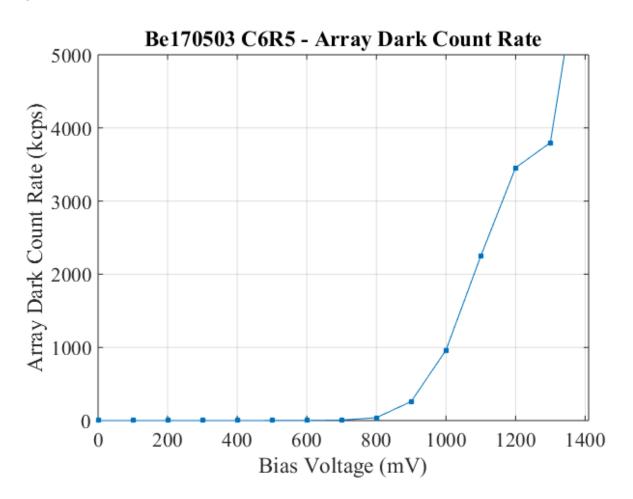
- Full system timing jitter is typically below 250 ps FWHM
- Dominated by jitter in FPGA-based time-to-digital converter
- Timing offsets between channels are due to imperfect length matching in readout chain





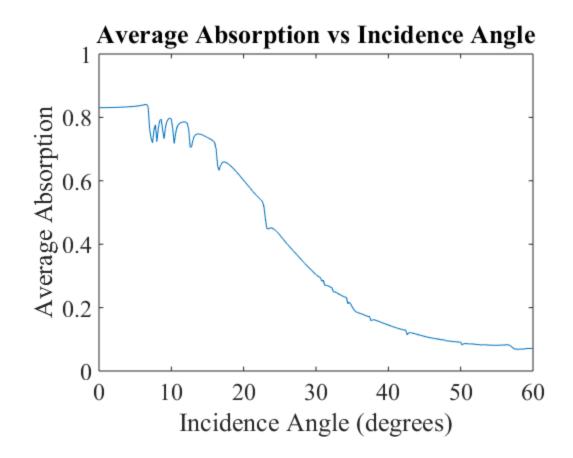


- 350 kcps at back of plateau 3.8 Mcps at front of plateau
- Can implement cryogenic spatial filter and/or shortpass filter to improve this
- Engineering tradeoff between false counts and efficiency





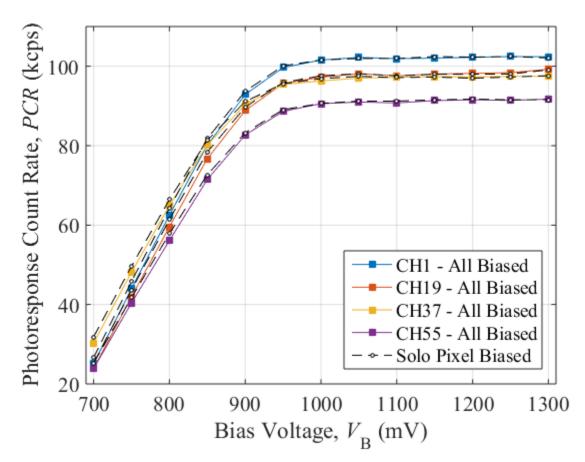
Angular Acceptance of Detector



- RCWA simulations predict that optical stack limits angular acceptance of detector
- 3 dB point is at a half-angle of 26 degrees
- Resonances are due to diffraction effects, break azimuthal symmetry



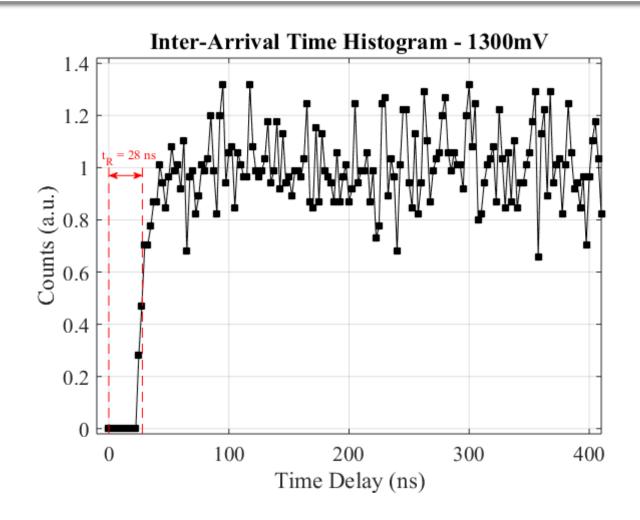
Absence of Crosstalk



- No crosstalk is observed with 1200 nm pitch co-wound arrays
- Severe crosstalk was observed with 320-800 nm pitch co-wound arrays
- From scaling, crosstalk is believed to be thermal
- Physics of crosstalk is under study with a generalized electrothermal model



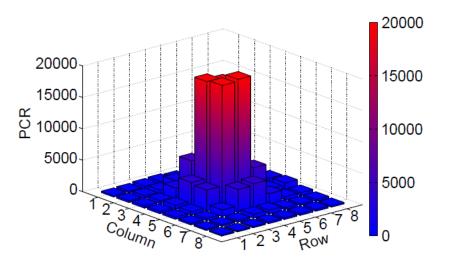
Absence of Afterpulsing



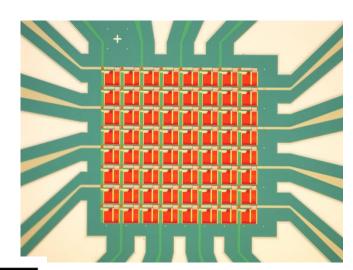
- Interarrival time histogram shows no presence of afterpulsing
- Recovery time is 28 ns

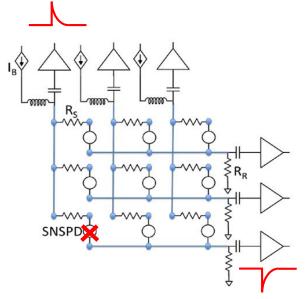


64 Pixel "Row-Column" Arrays



- 64 pixel (8 x 8) sparse WSi SNSPD array for fast time-correlated imaging
- Row-Column readout strategy allows 64 pixels to be read out using 16 lines
- Kilopixel Row-Column arrays are "lowhanging fruit" with 64-channel readout



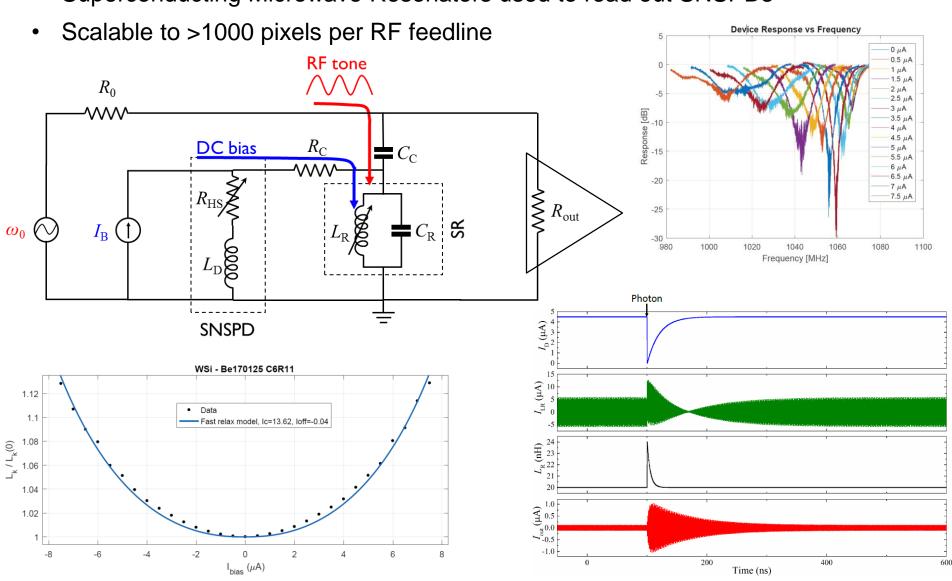


Operating Concept

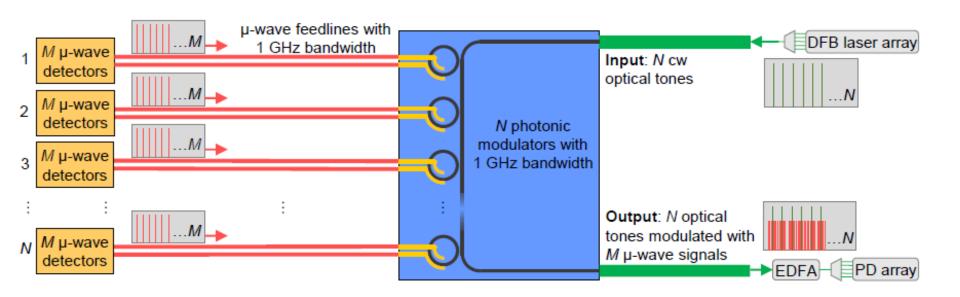


Frequency Multiplexed SNSPDs

Superconducting Microwave Resonators used to read out SNSPDs



Optical Multiplexing Strategies

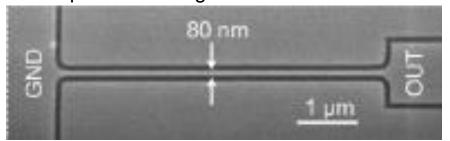


- Superconducting Microwave Resonators read out superconducting detectors
- Use cryogenic optical modulator to encode many "banks" of RF tones onto an optical carrier
- Scalable pathway to get megapixel-class superconducting camera data out of a sub-kelvin cryostat
- Especially relevant for Far-IR and deep UV astrophysics applications

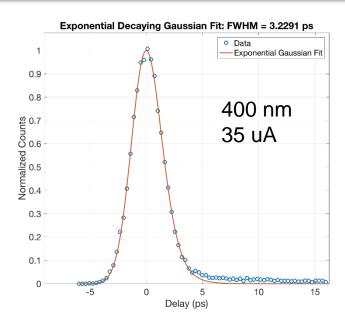


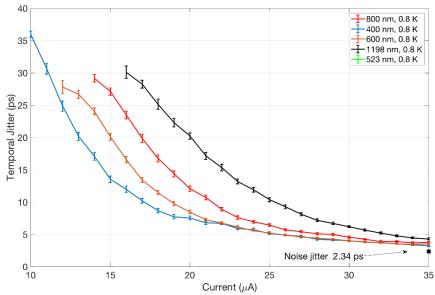
Ultra High Time Resolution SNSPDs

- Single photon detection with 3.2 ps FWHM time resolution achieved at 400 nm and 4.8 ps FWHM jitter at 1550 nm
- Specialized SNSPD with low-noise cryogenic amplifier readout
- 5 µm x 120 nm x 5 nm NbN nanowire
- Still limited by instrumental mechanisms at lowest jitter – have not yet reached fundamental limits
- Intrinsic jitter mechanism is now accessible for study
- Low jitter in devices with larger active area is practical using differential readout





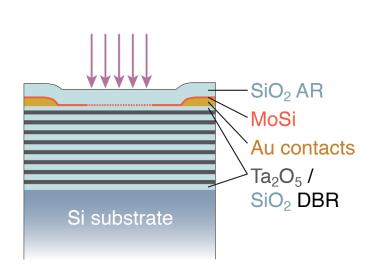


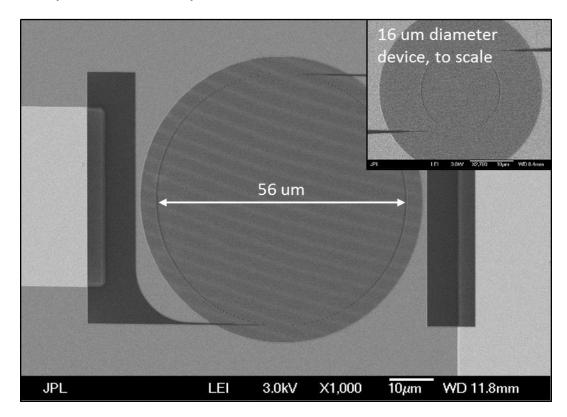




Ultraviolet SNSPDs

- Fiber-coupled MoSi UV SNSPDs for applications in ion trap quantum computing
- 80% Efficiency at 370 and 315 nm, single photon sensitivity at 245 nm
- DBR mirrors to enhance absorption
- 4.2 K operating temperature
- mHz dark count rates when coupled to optics, < 7e-5 cps intrinsic dark count rates

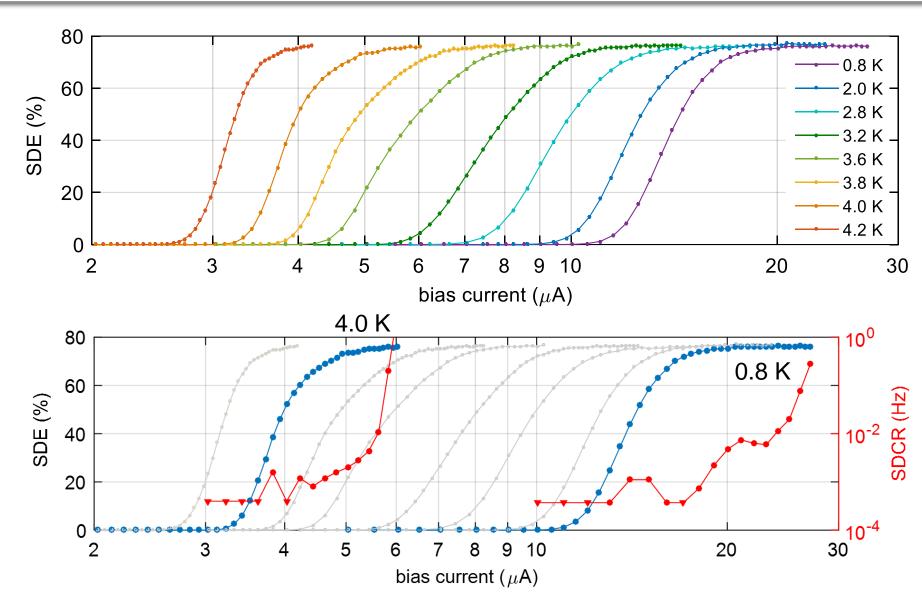








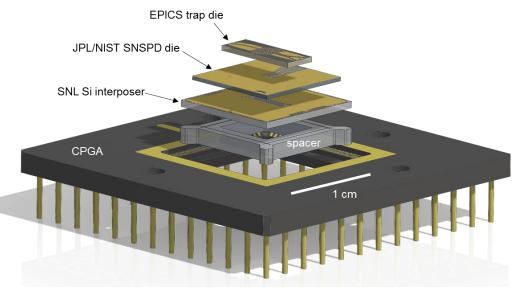
Efficiency and Dark Counts at 370nm



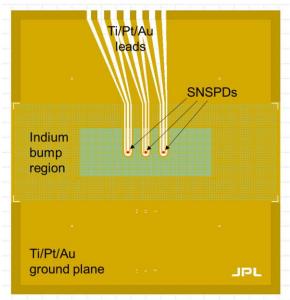
cavity

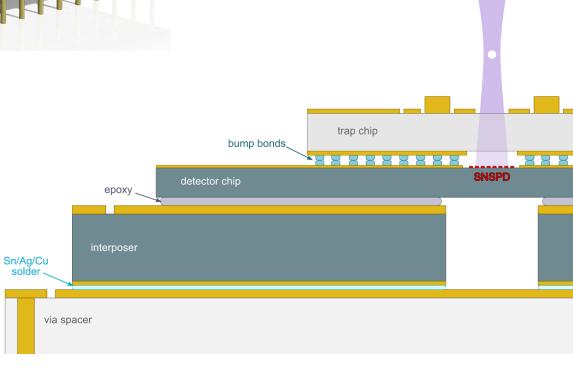


Integration with Ion Trap Chips



- Hybrid integration between ion trap chips and free-space UV SNSPDS
- Collaborative effort between JPL, NIST, Sandia, and Duke University

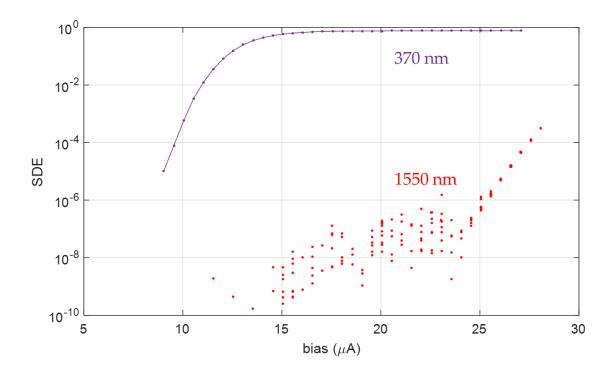






Intrinsic SNSPD Dark Counts

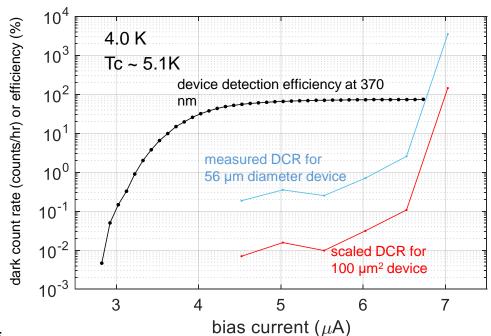
- SNSPDs optimized for high efficiency in the NIR are limited by thermal blackbody
- False count rate determined by spatial and spectral filtering
- UV SNSPDs are blind to IR, ideal for studying intrinsic false counts
- False counts <10⁻³ cps at 70% SDE coupled to fiber
- 56 µm diameter active area coupled to 50 µm core fiber

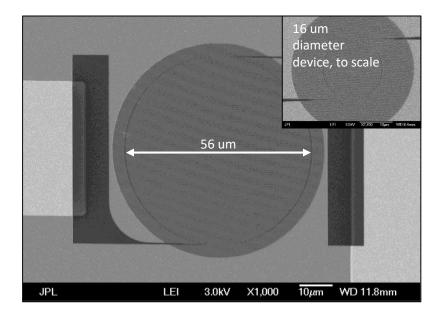




Intrinsic SNSPD Dark Counts

- To improve the statistics, returned to larger active area devices (56 µm diameter)
- Performed false count measurements at 4 K without coupling to fiber
- Each point integrated for 10 hours
- Measured 7e-5 cps (6 counts per day) on large-area device at 4 K
- If DCR scales with active area, 10 x 10 μm pixel would have 3e-6 cps (88 counts per year)

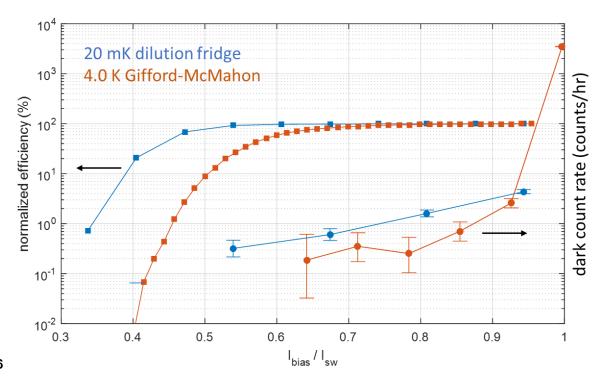






Intrinsic SNSPD Dark Counts

- Measured 56 µm device in a heavily shielded dilution refrigerator at 20 mK
- Illuminated with cryogenic UV LED at 4 K, integrated 28 hours per point
- False counts at 20 mK comparable to 4.0 K in less well shielded system
- Would expect temperature dependence for intrinsic dark counts

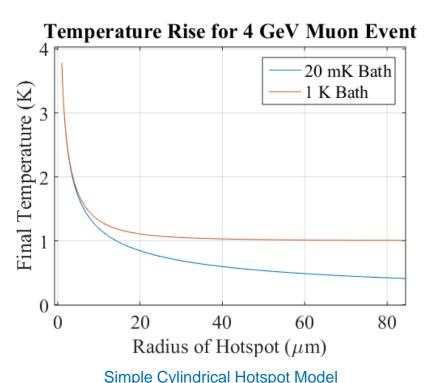


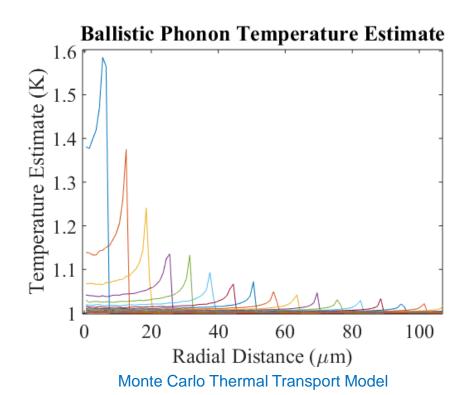




Predicted Cosmic Ray Limits

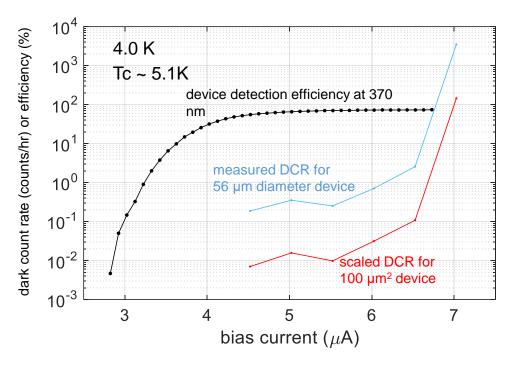
- At sea level, muons are the dominant contribution to cosmic ray flux
- Integrated flux rate 0.0148 muon / cm² / second, average energy 4 GeV
- Modeled thermal transport in Si, found < 10 µm sensitive region beyond active area
- For 60 µm diameter detectors needed to couple efficiently to 1-m telescope, muon count rate is 5.7e-8 cps – much lower than observed count rates

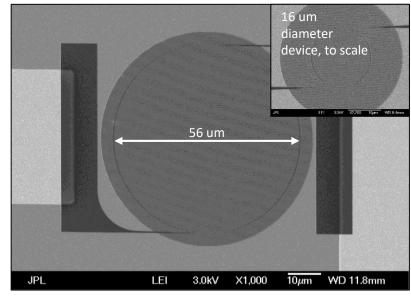






Ultra Low Dark Counts for UV SNSPDs



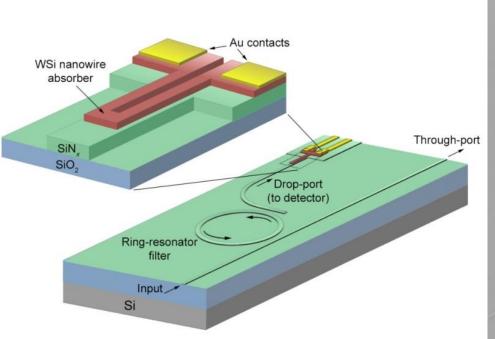


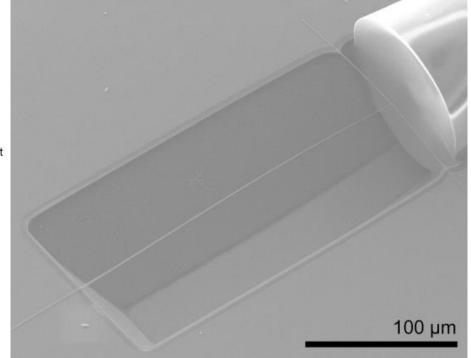
- 10 nm thick, 110 nm wide MoSi wire optimized for high efficiency (76%) at 370 nm
- UV SNSPDs have > 70 dB blindness to IR wavelengths, eliminating thermal background
- Measured 7e-5 cps (6 counts per day) on large-area device at 4 K
- If DCR scales with active area, 10 x 10 μm pixel would have 3e-6 cps (88 counts per year)



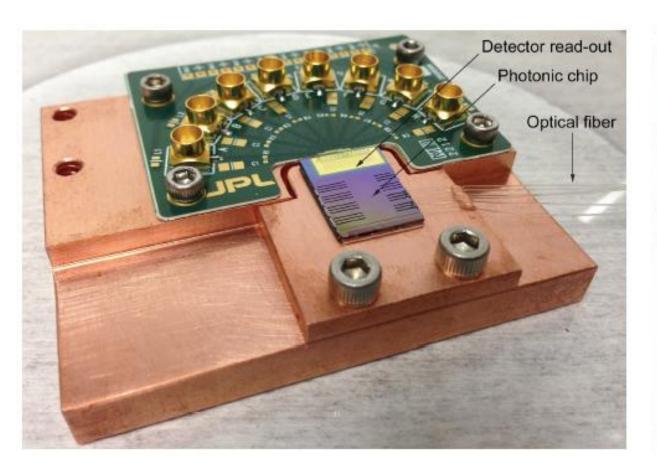
On-Chip Integrated SNSPDs

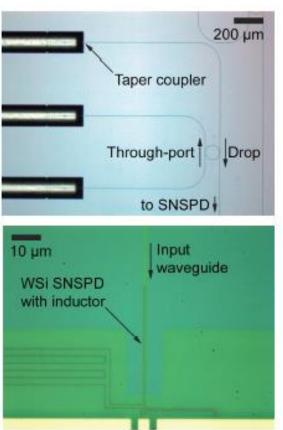
- WSi SNSPDs coupled to SiN waveguide photonics platform
- Integration with low-loss broadband optical couplers (Collaboration w/ Painter Group, Caltech)
- Integration with on-chip ring resonators or echelle grating to form channelizing spectrometer or DWDM receiver for QKD
- Can realize a robust, on-chip cryogenic spectrometer, particularly in the mid-IR
- Promising preliminary results



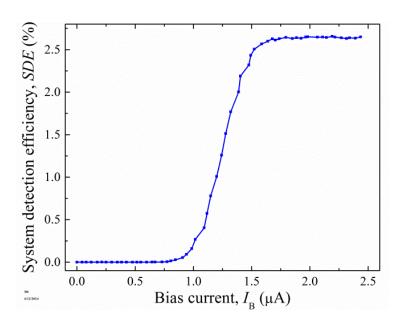


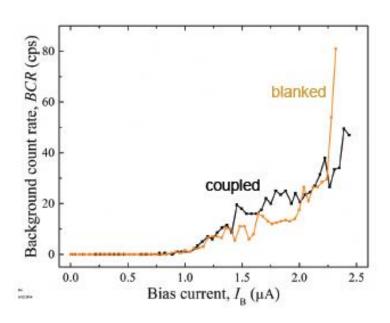


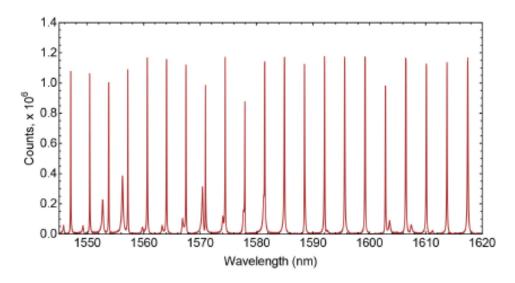


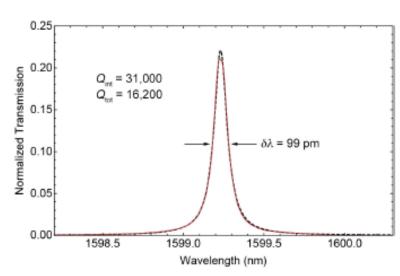


On-Chip Integrated SNSPDs









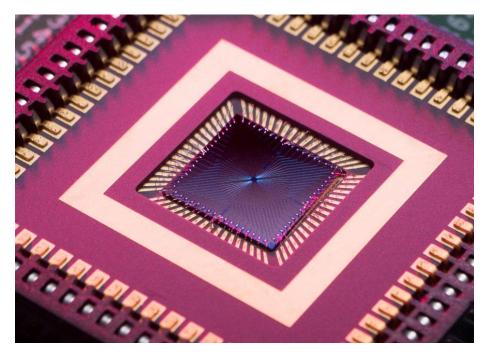
Future Directions for Waveguide SNSPDs

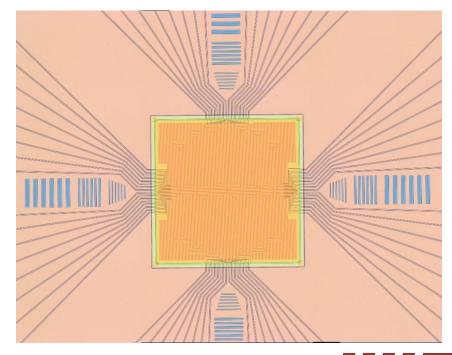
- Single-mode receivers behind AO systems
 - Integrated photonics is well suited to one or few spatial modes
- Ultra-low jitter optical communication receivers
 - Waveguide devices can be very short, reducing geometric jitter
- Frequency demultiplexing receivers
 - Integrated photonics is well suited to on-chip spectrometers, AWGs
- Ultra-high count rate receivers
 - Trees of low-loss beamsplitters can be used to feed light to SNSPDs
- Advanced receiver schemes
 - "Green Machine" can be implemented in photonic nanoprocessors
 - Gives information efficiency of PPM while alleviating peak-to-average power requirements on the transmit laser.
- On-chip heralded single photon sources
 - Relevant for space-to-ground quantum communication experiments



Conclusions

- SNSPDs are the highest performing detectors available for time correlated single photon counting
- Deep cryogenic operating temperatures make them best suited for the ground
- Progress in performance has been extremely rapid
- Technology is very new, with many opportunities for new innovation
- Fully compatible with integrated photonics







64-pixel SNSPD array mounted in chip carrier

